Research Article

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Heavy mesons mass spectroscopy under a spindependent Cornell potential within the framework of the spinless Salpeter equation

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Abstract: The energy bound-state solutions of the spinless Salpeter equation (SSE) have been obtained under a spin-dependent Cornell potential function via the Wentzel-Kramers-Brillouin approximation. The energy levels were applied to predict the mass spectra for the charmonium, bottomonium, and bottom-charmed mesons. The relativistic corrections for the angular momentum quantum number l > 0, total angular momentum quantum numbers j = l, $j = l \pm 1$, and the radial quantum numbers n = 1-4 improve the mass spectra. The results agree fairly with experimental data and theoretic results reported in existing works, where the authors utilized different forms of the inter-quark potentials and methods. The deviation of the obtained masses for the charmonium and bottomonium from the observed data yields a total percentage error of 3.32 and 1.11%, respectively. The results indicate that the accuracy of the masses is correlated with the magnitude of masses for the charm and bottom quarks. The SSE together with the phenomenological spin-dependent Cornell potential provides an adequate account of the mass spectroscopy for the heavy

mesons and may be used to predict other spectroscopic parameters.

Keywords: meson spectroscopy, spin-orbit coupling, Cornell potential, Salpeter equation

1 Introduction

The discovery of the electron by Thomson [1] in 1897 and the nucleus by Rutherford [2] in 1911 gave a fairly complete picture of the atomic structure. However, decades after these discoveries, several other elementary particles have been discovered with the aid of modern equipment such as particle colliders, detectors, and accelerators. These elementary particles that are the composite of much smaller particles are referred to as quarks. Quarks form the building blocks of matter, and their dynamical existence has been confirmed via experimental works on deep inelastic scattering of electrons and neutrons [3]. The standard model of elementary particles forms the basis for the understanding of particle physics, where the quarks have six flavors, namely, the up (u), down (d), top (t), bottom (b), strange (s), and charm (c) quarks; six leptons such as the electron (e), muon (μ), tau (τ), and their respective neutrinos (v_e , v_μ , v_τ); the gauge; and the Higgs bosons. The quarks and leptons interact via the unified electroweak forces and the strong quantum chromo-dynamics (QCD) force. The strong QCD force is responsible for the binding of quarks into protons and neutrons, while the interaction between the leptons is mediated by the electromagnetic force. The weak nuclear force is associated with particle decay.

The bound states of three quarks give rise to the baryons, while the quark-antiquark pairs constitute the mesons. Both the baryons and mesons are grouped into hadrons. The hadrons with ¹/₂ integer spin are fermions and obey the Fermi-Dirac statistics, while bosons possess spin 0, 1 and obey the Bose-Einstein statistics. The properties of these elementary particles have garnered research interest among particle physicists

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in recent years. These properties have been observed using heavy machinery [4], and some were predicted theoretically before their experimental discovery. In this regard, the boundstate solutions of the wave equations under the inter-quark potentials have been utilized to predict the mass spectroscopy and decay properties of the elementary particles. The most utilized potential energy is the Cornell potential, which is the combination of the Coulomb's energy and a linear function. The Coulomb's energy is responsible for the short-range gluon exchange interaction between a quark and its antiquark, while the linear function is in charge of quark confinement. The addition of spin components to the Cornell potential allows for relativistic corrections and results in the hyperfine splitting between the s-wave singlet and triplet states. The multiple triplet splitting occurs for any angular momentum quantum number l > 0.

Li et al. [5] predicted the charmonium ($c\bar{c}$) mass spectra using the coupled-channel model and the screened potential model in the mass region below 4 GeV. Their results agreed with the masses of the $c\bar{c}$ meson obtained with a quenched potential model and literature data [6]. Mutuk [7] investigated the mass spectra and decay constants of vector and pseudoscalar heavy-light mesons within the framework of the QCD sum rule and the quark model. The numerical results were in good agreement compared to observed data and other theoretical works. The charmonia spectra have been investigated using the non-relativistic quark model and matrix-Numerov method [8,9]. Chaturvedi and Rai [10], in a recent work, investigated the electromagnetic transitions, mass spectroscopy, and decay rates of the bottom-charm ($b\bar{c}$) meson within the context of the non-relativistic QCD. Several authors [11-28] have carried out extensive studies on the mesons bound-state solutions and applied them to obtain their spectroscopic parameters. The results in the references therein were compared to experimental data of the particle data group [29-31], and the results were predicted using different QCD-inspired potentials, phenomenological potentials, and theoretical methods.

In this work, we investigate the hyperfine mass spectra splitting of the heavy mesons within the framework of the semi-relativistic spinless Salpeter equation (SSE). Previously, the mass spectra of the heavy mesons have been obtained in previous studies [32,33] using the SSE without considering the spin components and relativistic corrections for the inter-quark potentials. The results in the references therein revealed that the semi-relativistic equation provides a satisfying account for the meson mass spectroscopy. Motivated by these facts, we report for the first time the approximate analytical and numerical mass spectra splitting of the heavy mesons under the SSE with a phenomenological spin-dependent Cornell potential *via* the Wentzel–Kramers–Brillouin (WKB) approximation.

In this study, we considered the interactions potential function given by

$$V(r) = V_{\rm c}(r) + V_{\rm SS}(r) + V_{\rm LS}(r) + V_{\rm T}(r),$$
(1)

where $V_{c}(r)$ is the Cornell potential. The functions $V_{SS}(r)$, $V_{LS}(r)$, and $V_{T}(r)$ are the spin–spin, spin–orbit, and tensor channels, respectively. The respective potential functions are represented as [22,24]

$$V_{\rm c}(r) = -\frac{4\alpha_s}{3r} + br, \qquad (2)$$

$$V_{\rm SS}(r) = \frac{32\pi\alpha_s}{9m_q m_{\bar{q}}} \bar{\delta}_\sigma(r) \langle \bar{\mathbf{S}} \cdot \bar{\mathbf{S}} \rangle, \tag{3}$$

$$V_{\rm LS}(r) = \frac{1}{m_q m_{\bar{q}}} \left(\frac{2\alpha_{\rm s}}{r^3} - \frac{b}{2r} \right) \langle \bar{L} \cdot \bar{S} \rangle, \tag{4}$$

$$V_{\rm T}(r) = \frac{4\alpha_s}{m_q m_{\bar{q}} r^3} \langle \bar{T} \rangle, \tag{5}$$

where α_s , b, m_q , and $m_{\bar{q}}$ are the coupling constant, linear confinement parameter, the mass of quark, and its antiquark, respectively. In (3), the Dirac delta function has been used as a Gaussian function $(\bar{\delta}_{\sigma}(r) = (\sigma/\sqrt{\pi})^3 e^{-\sigma^2 r^2})$. The operators in Eqs. (3)–(5) are diagonal in $|j, l, s\rangle$ with the respective spin–spin $((\bar{S} \cdot \bar{S}))$ spin–orbit $(\bar{L} \cdot \bar{S})$ and tensor (\bar{T}) matrix elements given by refs [22,24]

$$\langle \bar{\mathbf{S}} \cdot \bar{\mathbf{S}} \rangle = \frac{s(s+1)}{2} - \frac{3}{4},\tag{6}$$

$$\langle \bar{L} \cdot \bar{S} \rangle = \frac{1}{2} (j(j+1) - l(l+1) - s(s+1)),$$
 (7)

$$\langle n^3 l_j | \bar{\boldsymbol{T}} | n^3 l_j \rangle = -\frac{6(\langle \bar{\boldsymbol{L}} \cdot \bar{\boldsymbol{S}} \rangle)^2 + 3\langle \bar{\boldsymbol{L}} \cdot \bar{\boldsymbol{S}} \rangle - 2s(s+1)l(l+1)}{6(2l-1)(2l+3)}.$$
 (8)

The notations *s*, *l*, and j = l + s denote the spin number, orbital quantum number, and the total angular momentum quantum number, respectively. The spin–spin coupling gives rise to the s-wave (l = 0) hyperfine splitting between the triplet (s = 1) and singlet (s = 0) states. For l > 0, and $j = l \pm 1$, j = l, we have the multiplets splitting for the p, d, f, g, and h triplets quantum states. The n represents the principal quantum number. The spin-dependent potentials in (4) and (5) give the mass shifts and are obtained from leading-order perturbation theory [14,24]. The operator $\langle \bar{T} \rangle$ can be described by the non-vanishing diagonal matrix element for l > 0 and correspond to the spin triplet states [21].

2 Energy spectrum of the SSE with spin-dependent Cornell potential *via* the WKB approximation

To obtain an approximate analytical solution, we truncate the Gaussian function to a harmonic function for $r \ll 1$ fm *via* a Taylor series expansion around r = 0. The Gaussian function can be expressed as $e^{-\sigma^2 r^2} \sim 1 - \sigma^2 r^2$. In the femtometer scale, this approximation is important for quark interactions [28].

Using this approximation, the potential in (1) can be simplified as

$$V(r) = -\frac{Q}{r} + br + \frac{N}{r^{3}} - Pr^{2} + \Omega_{s},$$
 (9)

where

$$Q = \frac{4\alpha_s}{3} + \frac{b(\bar{L}\cdot\bar{S})}{2m_q m_{\bar{q}}},\tag{10}$$

$$N = \frac{2\alpha_s(\bar{L}\cdot\bar{S})}{m_a m_{\bar{a}}} + \frac{4\alpha_s(\bar{T})}{m_a m_{\bar{a}}},$$
(11)

$$P = \sigma^2 \Omega_s, \tag{12}$$

$$\Omega_s = \frac{16\pi\alpha_s}{9m_q m_q} (\sigma/\sqrt{\pi})^3 (s(s+1) - 1.5).$$
(13)

The spinless SSE equation for describing a two-body system is given as [34,35]

$$\left(\sum_{i=1,2}\sqrt{-\Delta+m_i^2}+V(r)-E_{nl}\right)\xi(r,\theta,\varphi)=0,\quad \Delta=\nabla^2,\quad (14)$$

where

$$\xi(r,\theta,\varphi) = \psi_{nl}(r)Y_{lm}(\theta,\varphi).$$
(15)

The notations ∇^2 , V(r), E_{nl} , and $\xi(r, \theta, \varphi)$ represent the Laplacian, the potential function, total energy, and the total wave function, respectively.

For interaction between particles, the summation in Eq. (14) can further be expanded *via* Taylor series to order two:

$$\sum_{i=1,2} \sqrt{-\Delta + m_i^2} = m_1 + m_2 - \frac{\Delta}{2\mu} - \frac{\Delta^2}{8\eta^3}, \quad (16)$$

where

$$\mu = \frac{m_1m_2}{m_1 + m_2}, \ \eta = \mu \left(\frac{m_1m_2}{m_1m_2 - 3\mu^2}\right)^{1/3}.$$

Eq. (14) can further be reduced to a Schrödinger-like equation [35]

$$\left(-\frac{1}{2\mu} \frac{d^2}{dr^2} + V(r) - E_{nl} - \frac{1}{2\tilde{m}} (V(r) - E_{nl})^2 + \frac{(l+1/2)^2}{2\mu r^2} \right) \psi_{nl}(r) = 0, \ \hbar = 1, \ \tilde{m} = \frac{\eta^3}{\mu^2}.$$

$$(17)$$

Eq. (17) can be written as a momentum eigenvalue equation

$$((\hat{\boldsymbol{P}})^2 + p_{\mu}^2(r))\psi_{nl}(r) = 0, \qquad (18)$$

where \hat{P} is the radial momentum operator and $p_{\mu}(r)$ is the meson momentum eigenvalue given as

 $p_{\mu}(r)$

$$= \sqrt{2\mu \left(\frac{1}{2\tilde{m}}(V(r) - E_{nl})^2 - (V(r) - E_{nl}) - \frac{(l+1/2)^2}{2\mu r^2}\right)}.$$
 (19)

To obtain the energy equation for the modeled potential, we employed the WKB energy quantization condition for two real turning points r_1 and r_2 via the integral equation:

$$\int_{r_1}^{r_2} p_{\mu}(r) dr = \pi \left(n + \frac{1}{2} \right), \hbar = 1.$$
 (20)

It is worth stating that we have added a Langer's correction [36] to the centrifugal potential using the transformation $l(l + 1) \rightarrow (l + 1/2)^2$. This correction in the WKB approximation admits the exact energy eigenvalues for soluble potentials and ensures that the wave function is well behaved near the origin.

Inserting the momentum into the WKB quantization integral in (20) with the potential energy given by (1), we obtained

$$\sqrt{2\mu} \int_{r_1}^{r_2} \sqrt{\frac{A}{r^6} - \frac{B}{r^4} + \frac{C}{r^3} + \frac{D}{r^2} - \frac{E}{r}} + \Lambda r^4 - Fr^3 + Gr^2 + Hr + K dr$$

$$= \pi \left(n + \frac{1}{2} \right),$$
(21)

where

$$A = \frac{N^2}{2\tilde{m}}, \quad B = \frac{2QN}{2\tilde{m}}, \quad C = \frac{2N\Omega_s - 2NE_{nl}}{2\tilde{m}} - N, \quad D = \frac{Q^2 + 2Nb}{2\tilde{m}} - L_s$$

$$E = \frac{2Q\Omega_s + 2PN - 2QE_{nl}}{2\tilde{m}} - Q, \quad F = \frac{2Pb}{2\tilde{m}}, \quad G = \frac{b^2 - 2P\Omega_s + 2PE_{nl}}{2\tilde{m}} + P_s$$

$$H = \frac{2b\Omega_s + 2PQ - 2bE_{nl}}{2\tilde{m}} - b, \quad K = \frac{(E_{nl} - \Omega_s)^2 - 2Qb}{2\tilde{m}} + E_{nl} - \Omega_s, \quad \Lambda = \frac{P^2}{2\tilde{m}}$$

$$L = \frac{(l + 1/2)^2}{2u}.$$

Using coordinate transformation q = 1/r, Eq. (21) reduces to

$$-\sqrt{2\mu} \int_{q_1}^{q_2} \sqrt{Aq^2 - B + \frac{C}{q} + \frac{D}{q^2} - \frac{E}{q^3} + \frac{\Lambda}{q^8} - \frac{F}{q^7} + \frac{G}{q^6} + \frac{H}{q^5} + \frac{K}{q^4}}$$

$$dq = \pi \left(n + \frac{1}{2}\right).$$
(22)

To solve Eq. (22) analytically, the multiple turning points need to be reduced to two *via* a Pekeris-type approximation around q = 0. Let $q = y + \delta$ with $\delta(1/q)$ assumed to be the characteristic distance between the quark and antiquark pairs. The inverse power terms can be obtained using the Taylor series expansion to the second order:

$$q^{-k} = \delta^{-k} \left(1 + \frac{y}{\delta} \right)^{-k} \sim \delta^{-k} - k \delta^{-k-1} y + \frac{k(k+1)}{2!} \delta^{-k-2} y^2 + O(y^3),$$
(23)

| Parameters | Potential free parameters | | | | | | |
|------------------------------|--|---------------------------------|----------------------------|--|--|--|--|
| | Charmonium | Bottomonium | Bc Meson | | | | |
| Mass (GeV) | $m_c = m_{\bar{c}} = 1.530^{\text{a}}$ | $m_b = m_{\bar{b}} = 5.250^{a}$ | $m_c = 1.530, m_b = 5.250$ | | | | |
| as | 30.1602 | 27.7862 | 41.1745 | | | | |
| δ (GeV) | 0.05846 | 0.0796 | 0.0796 | | | | |
| <i>b</i> (GeV ²) | 0.1416 | 0.1702 | 0.2710 | | | | |
| σ (GeV) | -0.1294 | 0.1356 | 0.1356 | | | | |

Table 1: Calculated parameters of interaction potential function

Note: ^aWe choose the bottom and charm masses from the obtained range [29] $1.2 < m_c < 1.8$ GeV and $4.5 < m_b < 5.4$ GeV.

Table 2: S and P-states mass spectrum of charmonium meson in GeV

| State | | Present | [24] | [18] | [16] | [11] | [22] | Expt. [30] |
|----------------------------|-----------------|---------|-------|--------|-------|--------|--------|----------------------------------|
| $n^{2s+1}L_j$ | J ^{PC} | | | | | | | |
| $J/\psi(1^{3}S_{1})$ | 1 | 3.520 | 3.094 | 3.0413 | 3.096 | 3.126 | 3.0851 | $3.097 \pm (6 \times 10^{-6})$ |
| $\eta_c(1^1S_0)$ | 0^+ | 3.293 | 2.989 | 3.1404 | 2.981 | 3.033 | 2.9904 | 2.984 ± 0.0004 |
| $\psi(2^{3}S_{1})$ | 1 | 3.902 | 3.681 | 3.7017 | 3.685 | 3.701 | 3.6821 | $3.686 \pm (6 \times 10^{-5})$ |
| $\eta_{c}(2^{1}S_{0})^{*}$ | 0-+ | 3.638 | 3.602 | 3.6610 | 3.635 | 3.666 | 3.6465 | $3.638 \pm (1.1 \times 10^{-3})$ |
| $\psi(3^3S_1)$ | 1 | 4.194 | 4.129 | 4.0502 | 4.039 | 4.055 | 4.1002 | 4.039 ± 10^{-3} |
| $\eta_c(3^1S_0)$ | 0^+ | 3.895 | 4.058 | 4.1347 | 3.989 | 4.158 | 4.0719 | |
| $\psi(4^{3}S_{1})^{*}$ | 1 | 4.421 | 4.514 | 4.4185 | 4.427 | 4.415 | 4.4394 | $4.421 \pm (4 \times 10^{-3})$ |
| $\eta_c(4^1S_0)$ | 0^+ | 4.091 | 4.448 | 4.4136 | 4.401 | 4.415 | 4.4209 | |
| $\psi(5^3S_1)$ | 1 | 4.600 | 4.863 | 4.6591 | 4.837 | 4.585 | | $4.63 \pm (6 \times 10^{-3})$ |
| $\eta_c(5^1S_0)$ | 0-+ | 4.242 | 4.799 | 4.6618 | 4.811 | 4.607 | | |
| $\psi(6^3S_1)$ | 1 | 5.804 | 5.185 | 4.8801 | 5.167 | 4.733 | | |
| $\eta_c(6^1S_0)$ | 0-+ | 5.175 | 5.124 | 4.8825 | 5.155 | 4.754 | | |
| $\chi_{c_1}(1^3P_1)^*$ | 1++ | 3.511 | 3.468 | 3.5036 | 3.511 | 3.487 | 3.5004 | $3.511 \pm (5 \times 10^{-5})$ |
| $\chi_{c_2}(1^3P_2)$ | 2++ | 3.545 | 3.480 | 3.4888 | 3.555 | 3.522 | 3.5514 | $3.556 \pm 7 \times 10^{-5}$ |
| $\chi_{c_0}(1^3 P_0)$ | 0++ | 3.466 | 3.428 | 3.4137 | 3.413 | 3.407 | 3.3519 | $3.415 \pm 3 \times 10^{-4}$ |
| $h_c(1^1 P_1)$ | 1+- | 3.298 | 3.470 | 3.5180 | 3.525 | 3.502 | 3.5146 | $3.525 \pm (1.1 \times 10^{-4})$ |
| $\chi_{c_1}(2^3P_1)$ | 1++ | 3.895 | 3.938 | 3.8072 | 3.906 | 3.786 | 3.9335 | $3.872 \pm (6 \times 10^{-5})$ |
| $\chi_{c_2}(2^3P_2)^*$ | 2++ | 3.923 | 3.955 | 3.9151 | 3.949 | 3.905 | 3.9798 | $3.923 \pm 1 \times 10^{-3}$ |
| $\chi_{c_0}(2^3 P_0)$ | 0++ | 3.856 | 3.897 | 3.7646 | 3.870 | 3.899 | 3.8357 | $3.922 \pm 1.8 \times 10^{-3}$ |
| $h_c(2^1 P_1)$ | 1+- | 3.642 | 3.943 | 3.8239 | 3.926 | 3.8210 | 3.9446 | |
| $\chi_{c_1}(3^3P_1)$ | 1++ | 4.188 | 4.338 | 4.1210 | 4.319 | 4.1230 | 4.3179 | $4.147 \pm (3 \times 10^{-3})$ |
| $\chi_{c_2}(3^3P_2)$ | 2++ | 4.212 | 4.358 | 4.1514 | 4.354 | 4.144 | 4.3834 | |
| $\chi_{co}(3^3 P_0)$ | 0++ | 4.155 | 4.296 | 4.0804 | 4.301 | 4.120 | 4.2167 | |
| $h_c(3^1 P_1)$ | 1+- | 3.899 | 4.344 | 4.1368 | 4.337 | 4.1640 | 4.3339 | |
| $\chi_{c_1}(4^3P_1)$ | 1++ | 4.416 | 4.696 | 4.4005 | 4.728 | 4.3730 | 4.6203 | |
| $\chi_{c_2}(4^3P_2)$ | 2++ | 4.436 | 4.718 | 4.4298 | 4.763 | 4.411 | 4.7367 | |
| $\chi_{c_0}(4^3P_0)$ | 0++ | 4.388 | 4.653 | 4.3621 | 4.698 | 4.362 | 4.5518 | |
| $h_c(4^1P_1)$ | 1+- | 4.094 | 4.704 | 4.1455 | 4.744 | 4.4200 | 4.6395 | |

where

$$\begin{split} q^{-1} &= \frac{3}{\delta} - \frac{3q}{\delta^2} + \frac{3q^2}{\delta^3}, \qquad q^{-2} = \frac{6}{\delta^2} - \frac{8q}{\delta^3} + \frac{3q^2}{\delta^4}, \qquad q^{-3} = \\ \frac{10}{\delta^3} - \frac{15q}{\delta^4} + \frac{6q^2}{\delta^5}, \quad q^{-4} = \frac{15}{\delta^4} - \frac{24q}{\delta^5} + \frac{10q^2}{\delta^6}, \quad q^{-5} = \frac{21}{\delta^5} - \frac{35q}{\delta^6} + \frac{15q^2}{\delta^7}, \\ q^{-6} &= \frac{28}{\delta^6} - \frac{48q}{\delta^7} + \frac{21q^2}{\delta^8}, \qquad q^{-7} = \frac{36}{\delta^7} - \frac{63q}{\delta^8} + \frac{28q^2}{\delta^9}, \qquad q^{-8} = \frac{45}{\delta^8} - \frac{80q}{\delta^9} + \frac{36q^2}{\delta^{10}}. \end{split}$$

Inserting the terms $q^{-k}(k = 1-8)$) into (22) with algebraic simplifications gives

$$-\sqrt{2\mu W} \int_{q_1}^{q_2} \sqrt{-q^2 + \varkappa q - T} \, \mathrm{d}q$$

$$= \sqrt{2\mu W} \int_{q_2}^{q_1} \sqrt{(q_2 - q)(q - q_1)} \, \mathrm{d}q = \pi \left(n + \frac{1}{2}\right),$$
(24)

where

$$\begin{split} & \kappa = \frac{1}{W} \left(\frac{63F}{\delta^8} - \frac{80\Lambda}{\delta^9} - \frac{48G}{\delta^7} - \frac{35H}{\delta^6} - \frac{24K}{\delta^5} + \frac{15E}{\delta^4} \right. \\ & \left. - \frac{8D}{\delta^3} - \frac{3C}{\delta^2} \right], \end{split} \tag{25}$$

$$T = \frac{1}{W} \left(\frac{36F}{\delta^7} - \frac{45\Lambda}{\delta^8} - \frac{28G}{\delta^6} - \frac{21H}{\delta^5} - \frac{15K}{\delta^4} + \frac{10E}{\delta^3} \right. \\ & \left. - \frac{6D}{\delta^2} - \frac{3C}{\delta} + B \right], \end{aligned} \tag{26}$$

$$W = \frac{28F}{\delta^9} - \frac{36\Lambda}{\delta^{10}} - \frac{21G}{\delta^8} - \frac{15H}{\delta^7} - \frac{10K}{\delta^6} + \frac{6E}{\delta^5} - \frac{3D}{\delta^4} \\ & \left. - \frac{3C}{\delta^3} - A. \right. \end{aligned}$$

The turning points (q_1, q_2) in (24) are obtained by solving the $\sqrt{-q^2 + \varkappa q - T} = 0$, which is a requirement in the WKB approximation for the momentum to vanish at the classical turning points:

| Table 3: D, F, and G-sta | tes mass spectrum | of charmonium | meson in GeV |
|--------------------------|-------------------|---------------|--------------|
|--------------------------|-------------------|---------------|--------------|

| State | | Present | [24] | [18] | [16] | [11] | [22] | Expt. [30] |
|-------------------------------|-----------------|---------|-------|---------|-------|--------|--------|--------------------------------|
| $n^{2s+1}L_j$ | J ^{PC} | | | | | | | |
| $\psi_2(1^3D_2)$ | 2 | 3.520 | 3.772 | 3.46047 | 3.795 | 3.3480 | 3.8077 | |
| $\psi_3(1^3D_3)$ | 3 | 3.575 | 3.755 | 3.51402 | 3.813 | 3.307 | 3.8146 | |
| $\psi_1(1^3D_1)$ | 1 | 3.461 | 3.775 | 3.40228 | 3.783 | 3.374 | 3.7853 | $3.774 \pm 4 \times 10^{-4}$ |
| $\eta_{c_2}(1^1D_2)$ | 2-+ | 3.306 | 3.765 | 3.47795 | 3.807 | 3.3760 | 3.8073 | |
| $\psi_2(2^3D_2)$ | 2 | 3.902 | 4.188 | 3.81161 | 4.190 | 3.8010 | 4.1737 | $3.824 \pm (5 \times 10^{-4})$ |
| $\psi_3(2^3D_3)$ | 3 | 3.949 | 4.176 | 3.86300 | 4.220 | 3.797 | 4.1829 | |
| $\psi_1(2^3D_1)$ | 1 | 3.852 | 4.188 | 3.75606 | 4.150 | 3.800 | 4.1504 | |
| $\eta_{c_2}(2^1\mathrm{D}_2)$ | 2-+ | 3.649 | 4.182 | 3.82825 | 4.196 | 3.8360 | 4.1737 | |
| $\psi_2(3^3D_2)$ | 2 | 4.195 | 4.557 | 4.12502 | 4.544 | 4.1350 | 4.5588 | |
| $\psi_3(3^3D_3)$ | 3 | 4.234 | 4.549 | 4.17431 | 4.574 | 4.163 | 4.5725 | |
| $\psi_1(3^3D_1)$ | 1 | 4.152 | 4.555 | 4.07208 | 4.507 | 4.113 | 4.5258 | |
| $\eta_{c_2}(3^1D_2)$ | 2-+ | 3.905 | 4.553 | 4.14085 | 4.549 | 4.176 | 4.5597 | |
| $\chi_{c_3}(1^3\mathrm{F}_3)$ | 3++ | 3.532 | 4.012 | 3.46746 | 4.068 | 3.375 | 4.0440 | |
| $\chi_{c_4}(1^3F_4)$ | 4++ | 3.609 | 4.036 | 3.54185 | 4.093 | 3.315 | 4.0374 | |
| $\chi_{c_2}(1^3 F_2)$ | 2++ | 3.454 | 3.990 | 3.38961 | 4.041 | 3.403 | 4.0424 | |
| $\eta_{c_3}(1^1F_3)$ | 3+- | 3.319 | 4.017 | 3.48494 | 4.071 | 3.403 | 4.0411 | |
| $\chi_{c_3}(2^3F_3)$ | 3++ | 3.913 | 4.396 | 3.81815 | 4.400 | 3.823 | 4.3744 | |
| $\chi_{c_4}(2^3F_4)$ | 4++ | 3.978 | 4.415 | 3.88949 | 4.434 | 3.814 | 4.3711 | |
| $\chi_{c_2}(2^3F_2)$ | 2++ | 3.846 | 4.378 | 3.74376 | 4.361 | 3.812 | 4.3699 | |
| $\eta_{c_3}(2^1F_3)$ | 3+- | 3.660 | 4.400 | 3.83479 | 4.406 | 3.858 | 4.3723 | |
| $\psi_{c_4}(1^3G_4)$ | 4 | 3.549 | | | 4.343 | | 4.2506 | |
| $\psi_{c_{5}}(1^{3}G_{5})$ | 5 | 3.647 | | | 4.357 | | 4.2369 | |
| $\psi_{c_3}(1^3G_3)$ | 3 | 3.450 | | | 4.321 | | 4.2582 | |
| $\eta_{c_4}(1^1G_4)$ | 4-+ | 3.336 | | | 4.345 | | 4.2471 | |
| Total error (χ) | | 3.32% | 1.98% | 1.64% | 1.30% | 1.33% | 1.53% | |

| Table 4: S ar | nd P-states | mass spe | ctrum of | bottomonium | meson in (| ъV |
|---------------|-------------|----------|-----------|---------------|------------|----|
| | ia i states | mass spe | cu ann or | Socconnonnann | meson me | |

| State | | Present | [17] | [16] | [18] | [11] | [14] | [24] | Expt. [30] |
|-------------------------|----------|---------|--------|--------|----------|--------|---------|---------|--|
| $n^{2s+1}L_j$ | J^{PC} | | | | | | | | |
| $Y(1^{3}S_{1})$ | 1 | 9.906 | 9.465 | 9.460 | 9.49081 | 9.525 | 9.4600 | 9.4600 | $9.460 \pm (2.6 \times 10^{-4})$ |
| $\eta_b(1^1S_0)$ | 0-+ | 9.916 | 9.402 | 9.398 | 9.43601 | 9.472 | 9.3900 | 9.4280 | $9.399 \pm (2 \times 10^{-3})$ |
| $Y(2^{3}S_{1})$ | 1 | 10.240 | 10.003 | 10.023 | 10.01257 | 10.049 | 10.0150 | 9.9790 | $10.023 \pm (3.1 \times 10^{-4})$ |
| $\eta_b(2^1S_0)$ | 0-+ | 10.251 | 9.976 | 9.990 | 9.99146 | 10.028 | 9.9900 | 9.9550 | |
| $Y(3^{3}S_{1})$ | 1 | 10.504 | 10.354 | 10.355 | 10.32775 | 10.371 | 10.3430 | 10.3590 | $10.355 \pm (5 \times 10^{-4})$ |
| $\eta_b(3^1S_0)$ | 0-+ | 10.517 | 10.336 | 10.329 | 10.1386 | 10.360 | 10.3260 | 10.3380 | |
| $Y(4^{3}S_{1})$ | 1 | 10.715 | 10.635 | 10.586 | 10.5461 | 10.598 | 10.5970 | 10.6830 | $10.579 \pm (1.2 \times 10^{-3})$ |
| $\eta_b(4^1S_0)$ | 0-+ | 10.729 | 10.623 | 10.573 | 10.3236 | 10.592 | 10.5840 | 10.6630 | |
| $Y(5^{3}S_{1})$ | 1 | 10.883 | 10.878 | 10.851 | 10.82628 | 10.870 | 10.8110 | 10.9750 | $10.885 \pm (2.6 \times 10^{-3}) \pm (1.6 \times 10^{-3})$ |
| $\eta_b(5^1S_0)$ | 0-+ | 10.899 | 10.869 | 10.869 | 10.4977 | 10.790 | 10.8000 | 10.9560 | |
| $Y(6^{3}S_{1})^{*}$ | 1 | 11.020 | 11.102 | 11.061 | 10.97061 | 11.022 | 10.9970 | 11.2430 | $11.020 \pm (4 \times 10^{-3})$ |
| $\eta_b(6^1S_0)$ | 0-+ | 11.036 | 11.097 | 11.088 | 10.6615 | 10.961 | 10.9880 | 11.2260 | 11.014 [17] |
| $\chi_{b_1}(1^3P_1)$ | 1++ | 9.906 | 9.876 | 9.892 | 9.87371 | 9.875 | 9.9030 | 9.8190 | $9.893 \pm (2.6 \times 10^{-4}) \pm (3.1 \times 10^{-4})$ |
| $\chi_{b_2}(1^3P_2)^*$ | 2++ | 9.912 | 9.897 | 9.912 | 9.89083 | 9.903 | 9.921 | 9.825 | $9.912 \pm (2.6 \times 10^{-4}) \pm (3.1 \times 10^{-4})$ |
| $\chi_{b_0}(1^3 P_0)$ | 0++ | 9.898 | 9.847 | 9.859 | 9.8432 | 9.840 | 9.864 | 9.806 | $9.859 \pm (4.2 \times 10^{-4}) \pm (3.1 \times 10^{-4})$ |
| $h_b(1^1P_1)$ | 1+- | 9.918 | 9.882 | 9.900 | 9.87919 | 9.884 | 9.9090 | 9.8210 | $9.899 \pm (8 \times 10^{-4})$ |
| $\chi_{b_1}(2^3 P_1)$ | 1++ | 10.240 | 10.246 | 10.255 | 10.21695 | 10.229 | 10.249 | 10.2170 | $10.255 \pm (2.2 \times 10^{-4}) \pm (5 \times 10^{-4})$ |
| $\chi_{b_2}(2^3 P_2)$ | 2++ | 10.245 | 10.261 | 10.268 | 10.22961 | 10.254 | 10.246 | 10.224 | $10.269 \pm (2.2 \times 10^{-4}) \pm (5 \times 10^{-4})$ |
| $\chi_{b_0}(2^3 P_0)^*$ | 0++ | 10.233 | 10.226 | 10.233 | 10.19625 | 10.202 | 10.220 | 10.205 | $10.233 \pm (4 \times 10^{-4}) \pm (5 \times 10^{-4})$ |
| $h_b(2^1P_1)$ | 1+- | 10.254 | 10.250 | 10.260 | 10.22153 | 10.237 | 10.254 | 10.220 | $10.260 \pm (1.2 \times 10^{-3})$ |
| $\chi_{h_1}(3^3P_1)$ | 1++ | 10.504 | 10.538 | 10.541 | 10.1378 | 10.339 | 10.515 | 10.553 | $10.514 \pm (7 \times 10^{-4})$ |
| $\chi_{h_2}(3^3 P_2)$ | 2++ | 10.509 | 10.550 | 10.550 | 10.1405 | 10.406 | 10.528 | 10.560 | $10.524 \pm (8 \times 10^{-4})$ |
| $\chi_{b_0}(3^3 P_0)$ | 0++ | 10.498 | 10.522 | 10.521 | 10.1342 | 10.299 | 10.490 | 10.540 | 10.500 [17] |
| $h_b(3^1P_1)^*$ | 1+- | 10.519 | 10.541 | 10.544 | 4.14695 | 10.362 | 10.5190 | 10.556 | 10.519 [17] |
| $\chi_{b_1}(4^3P_1)$ | 1++ | 10.715 | 10.788 | 10.802 | 10.3229 | 10.571 | | 10.853 | |
| $\chi_{b_2}(4^{3}P_2)$ | 2++ | 10.719 | 10.798 | 10.812 | 10.3255 | 10.637 | | 10.860 | |
| $\chi_{b_0}(4^3 P_0)$ | 0++ | 10.710 | 10.775 | 10.781 | 10.3193 | 10.532 | | 10.840 | |
| $h_b(4^1P_1)$ | 1+- | 10.731 | 10.790 | 10.804 | 10.3242 | 10.594 | | 10.855 | |

$$q_{1} = \frac{\varkappa}{2} - \frac{1}{2}\sqrt{\varkappa^{2} - 4T},$$

$$q_{2} = \frac{\varkappa}{2} + \frac{1}{2}\sqrt{\varkappa^{2} - 4T}.$$
(28)
(29)

Solving (24), we obtained

$$\int_{q_2}^{q_1} \sqrt{(q_2 - q)(q - q_1)} \, \mathrm{d}q = \frac{\pi}{8} (q_1 - q_2)^2.$$
(30)

Comparing Eqs. (24) and (30), the condition for the energy-level equation is obtained as

$$\frac{\varkappa^2}{4} - T = \sqrt{\frac{2}{\mu W}} \left(n + \frac{1}{2} \right).$$
(31)

The mass spectra are obtained from the relation between the quark masses and the energy eigenvalue:

$$M_{nl} = m_q + m_{\bar{q}} + E_{nl}.$$
 (32)

3 Numerical results and discussion

The energy bound-state solution of the SSE under a spindependent Cornell potential has been obtained *via* the semiclassical WKB approximation method. The potential parameters (α_s , b, δ , σ) were obtained by fitting the obtained

| $n^{2s+1}L_j$ | J^{PC} | Present | [17] | [16] | [18] | [11] | [14] | [24] | Expt. [30] |
|-------------------------------|----------|---------|--------|--------|---------|--------|--------|--------|-----------------------------------|
| $Y_2(1^3D_2)$ | 2 | 9.911 | 10.147 | 10.161 | 10.1126 | 10.096 | 10.153 | 10.075 | $10.164 \pm (1.4 \times 10^{-3})$ |
| $Y_3(1^3D_3)$ | 3 | 9.921 | 10.155 | 10.166 | 9.73855 | 9.849 | 10.157 | 10.073 | 10.172 [17] |
| $Y_1(1^3D_1)$ | 1 | 9.900 | 10.138 | 10.154 | 9.72905 | 9.666 | 10.146 | 10.074 | 10.155 [17] |
| $\eta_{b_2}(1^1 D_2)$ | 2-+ | 9.923 | 10.148 | 10.163 | 9.7355 | 9.767 | 10.153 | 10.074 | 10.165 [17] |
| $Y_2(2^3D_2)$ | 2 | 10.244 | 10.449 | 10.443 | 9.94259 | 10.071 | 10.432 | 10.424 | |
| $Y_3(2^3D_3)$ | 3 | 10.253 | 10.455 | 10.449 | 9.94704 | 10.175 | 10.436 | 10.423 | |
| $Y_1(2^3D_1)$ | 1 | 10.235 | 10.441 | 10.435 | 9.93775 | 9.996 | 10.425 | 10.423 | |
| $\eta_{b_2}(2^1\mathrm{D}_2)$ | 2-+ | 10.258 | 10.450 | 10.445 | 9.94405 | 10.093 | 10.432 | 10.424 | |
| $Y_2(3^3D_2)$ | 2 | 10.508 | 10.705 | 10.711 | 10.1391 | 10.345 | | 10.733 | |
| $Y_3(3^3D_3)$ | 3 | 10.516 | 10.711 | 10.717 | 10.1435 | 10.446 | | 10.733 | |
| $Y_1(3^3D_1)$ | 1 | 10.501 | 10.698 | 10.704 | 10.1344 | 10.272 | | 10.731 | |
| $\eta_{b_2}(3^1D_2)$ | 2-+ | 10.523 | 10.706 | 10.713 | 10.1405 | 10.368 | | 10.733 | |
| $\chi_{b_3}(1^3\mathrm{F}_3)$ | 3++ | 9.918 | 10.355 | 10.346 | 9.7361 | 9.754 | 10.340 | 10.287 | |
| $\chi_{b_4}(1^3\mathrm{F}_4)$ | 4++ | 9.932 | 10.358 | 10.349 | 9.74242 | 9.896 | 10.340 | 10.291 | |
| $\chi_{b_2}(1^3\mathrm{F}_2)$ | 2++ | 9.904 | 10.350 | 10.343 | 9.72948 | 9.642 | 10.338 | 10.283 | |
| $\eta_{b_3}(1^1F_3)$ | 3+- | 9.931 | 10.355 | 10.347 | 9.73759 | 9.778 | 10.339 | 10.288 | |
| $\chi_{b_3}(2^3\mathrm{F}_3)$ | 3++ | 10.251 | 10.619 | 10.614 | 9.94462 | 10.081 | | 10.607 | |
| $\chi_{b_4}(2^3\mathrm{F}_4)$ | 4++ | 10.263 | 10.622 | 10.617 | 9.9508 | 10.219 | | 10.609 | |
| $\chi_{b_2}(2^3F_2)$ | 2++ | 10.239 | 10.615 | 10.610 | 9.93815 | 9.971 | | 10.604 | |
| $\eta_{b_3}(2^1F_3)$ | 3+- | 10.265 | 10.619 | 10.615 | 9.94608 | 10.104 | | 10.607 | |
| $Y_4(1^3G_4)$ | 4 | 9.929 | 10.531 | 10.512 | | | | | |
| $Y_5(1^3G_5)$ | 5 | 9.946 | 10.532 | 10.514 | | | | | |
| $Y_3(1^3G_3)$ | 3 | 9.911 | 10.529 | 10.511 | | | | | |
| $\eta_{b_4}(1^1G_4)$ | 4-+ | 9.941 | 10.530 | 10.513 | | | | | |
| Total error (χ) | | 1.11% | 0.20% | 0.11% | 3.85% | 0.96% | 0.13% | 0.66% | |

Table 5: D, F, and G-states mass spectrum of bottomonium meson in GeV

mass spectra with the corresponding experimental masses of the particle data group in Tables 2–5. We present the parameters in Table 1. The quantum states in the asterisk correspond to the points where the potential parameters are fitted using (32). To check for accuracy, the potential parameters were obtained for random quantum numbers for the s-wave singlet and triplet states and also the hyperfine triplet states $(l > 0, j \ge 0)$ where we choose the parameters that reproduce minimum total errors. Also, to obtain potential parameters for the bottom-charmed meson, we assumed a constant characteristic distance (δ) and spin-dependent constant (σ) due to the limited availability of experimental data. For the bottomonium and charmonium mesons, we solved four polynomial equations simultaneously with the help of Maple software, while two non-linear equations were solved for the bottom-charmed meson to obtain the parameters a_s and *b*. The variation of the potential function with distance for the heavy mesons is plotted in Figure 1(a-c). The potential curves for the triplet and singlet states account for linear confinement and short-range gluon exchange between the

quark-antiquark pairs. The Coulomb part of the potential in the absence of tensor and spin orbit components dominates at short distances, whereas the linear component is prominent at large distances.

The total errors are obtained using the following formula:

$$\chi = \frac{100}{Z} \sum_{i=1}^{Z} \left| \frac{M_{nl}^{\exp} - M_{nl}^{\text{theo}}}{M_{nl}^{\exp}} \right|,$$
(33)

where Z, M_{nl}^{exp} , and M_{nl}^{theo} are the respective number of experimental data points, experimental masses, and theoretically obtained masses. In Table 2, the mass spectra of charmonium for the s-wave increase as the quantum number increases with the singlet states bounded below the triplet states. The $I/\psi(1^{3}S_{1})$ and $\eta_{c}(1^{1}S_{0})$ are higher than the experimentally determined values. However, the other s-wave masses are in agreement with the experimental masses [30] and the masses obtained using other interquark potential functions and methods reported in the existing literature [11,16,18,22,24]. In comparison with



Figure 1: (a–c) Variation of potential function with distance: (a) charmonium, (b) bottomonium, and (c) bottom-charm mesons for spin numbers s = 0 and s = 1.

experimental data, the masses $\chi_{c_2}(1^3P_2)$, $\chi_{c_0}(1^3P_0)$, $\psi_1(1^3D_1)$, and $\psi_2(2^3D_2)$ with hyperfine splitting and relativistic corrections l > 0, j = l, $j = l \pm 1$ and s = 1 as shown in Tables 2 and 3 were found to be more accurate compared to the masses ($h_c(1^1P_1)$, $\eta_{c_2}(1^1D_2)$, and $\eta_{c_2}(2^1D_2)$) for the quantum states (l > 0, s = 0, j = l). The charmonium masses for the s, p, d, and f states increase with the increase in the radial quantum number and were found to deviate from the experimental masses by a total error of 3.32%.

In Table 4, the bottomonium masses for the *s* and *p*-quantum states are presented. The low-lying quantum states masses are higher than the observed values. As the quantum number increases, the masses $Y(5^{3}S_{1})$ and $\eta_{b}(6^{1}S_{0})$ were found to be in good agreement compared to other works in the existing literature [11,14,16–18] and

the masses of the particle data group [30]. The *p*-states masses were found to agree with the ones obtained earlier. However, the masses for *d* quantum states presented in Table 5 fairly compare with the observed masses [30]. Generally, the masses increase with the increase in the radial quantum number. Using Eq. (33), the bottomonium mass spectra deviation from experimental data yields a total percentage error of approximately 1.11%, which indicates an improvement over the results in the study by Mansour and Gamal [18] and comparable to the total percentage error of 0.96% in the study by Mansour and Gamal [11] and 0.66% in the study by Soni *et al.* [24]. It can be seen that the percentage error is higher for the charmonium meson due to its light reduced mass. In Tables 6 and 7, the masses obtained for the bottom-charmed mesons

State

Table 6: S and P states mass spectrum of bottom-charmed meson in GeV

[16]

[11]

[24]

Expt. [31]

Present

| in GeV | | | | | | | | | | |
|-------------------------------|-----------------|---------|-------|-------|-------|--|--|--|--|--|
| $n^{2s+1}L_j$ | J ^{PC} | Present | [16] | [11] | [24] | | | | | |
| 1 ³ D ₂ | 2 | 6.228 | 7.025 | 6.299 | 6.997 | | | | | |

| $n^{2s+1}L_j$ | J^{PC} | | | | | |
|-----------------------------------|----------|-------|-------|-------|-------|-------|
| 1 ³ S ₁ | 1 | 6.225 | 6.333 | 6.313 | 6.321 | |
| (1 ¹ S ₀)* | 0-+ | 6.275 | 6.272 | 6.276 | 6.272 | 6.275 |
| $2^{3}S_{1}$ | 1 | 6.783 | 6.882 | 6.867 | 6.900 | |
| $(2^{1}S_{0})^{*}$ | 0-+ | 6.842 | 6.842 | 6.841 | 6.864 | 6.842 |
| $3^{3}S_{1}$ | 1 | 7.216 | 7.258 | 7.308 | 7.338 | |
| $3^{1}S_{0}$ | 0^+ | 7.283 | 7.226 | 7.281 | 7.306 | |
| $4^{3}S_{1}$ | 1 | 7.556 | 7.609 | 7.660 | 7.714 | |
| $4^{1}S_{0}$ | 0^+ | 7.631 | 7.585 | 7.634 | 7.684 | |
| $5^{3}S_{1}$ | 1 | 7.826 | 7.947 | 7.941 | 8.054 | |
| $5^{1}S_{0}$ | 0^+ | 7.907 | 7.928 | 7.917 | 8.025 | |
| $6^{3}S_{1}$ | 1 | 8.042 | | 8.168 | 8.368 | |
| $6^{1}S_{0}$ | 0-+ | 8.130 | | 8.144 | 8.340 | |
| $1^{3}P_{1}$ | 1++ | 6.218 | 6.743 | 6.281 | 6.705 | |
| $1^{3}P_{2}$ | 2++ | 6.248 | 6.761 | 6.366 | 6.712 | |
| $1^{3}P_{0}$ | 0++ | 6.178 | 6.699 | 6.223 | 6.686 | |
| $1^{1}P_{1}$ | 1+- | 6.280 | 6.750 | 6.290 | 6.706 | |
| $2^{3}P_{1}$ | 1++ | 6.777 | 7.134 | 6.836 | 7.165 | |
| $2^{3}P_{2}$ | 2++ | 6.803 | 7.157 | 6.917 | 7.173 | |
| $2^{3}P_{0}$ | 0++ | 6.743 | 7.146 | 6.782 | 7.146 | |
| $2^{1}P_{1}$ | 1+- | 6.847 | 7.147 | 6.846 | 7.168 | |
| $3^{3}P_{1}$ | 1++ | 7.211 | 7.500 | 7.278 | 7.555 | |
| $3^{3}P_{2}$ | 2++ | 7.233 | 7.524 | 7.355 | 7.565 | |
| $3^{3}P_{0}$ | 0++ | 7.182 | 7.474 | 7.227 | 7.536 | |
| $3^{1}P_{1}$ | 1+- | 7.287 | 7.510 | 7.287 | 7.559 | |
| $4^{3}P_{1}$ | 1++ | 7.552 | 7.844 | 7.631 | 7.905 | |
| $4^{3}P_{2}$ | 2++ | 7.571 | 7.867 | 7.704 | 7.915 | |
| $4^{3}P_{0}$ | 0++ | 7.527 | 7.817 | 7.583 | 7.885 | |
| $4^{1}P_{1}$ | 1+- | 7.634 | 7.853 | 7.640 | 7.908 | |

| $n^{2s+1}L_j$ | $J^{\mu c}$ | Present | [16] | [11] | [24] |
|-------------------------------|-------------|---------|-------|-------|-------|
| $1^{3}D_{2}$ | 2 | 6.228 | 7.025 | 6.299 | 6.997 |
| $1^{3}D_{3}$ | 3 | 6.278 | 7.029 | 6.429 | 6.990 |
| $1^{3}D_{1}$ | 1 | 6.177 | 7.021 | 6.200 | 6.998 |
| l^1D_2 | 2-+ | 6.291 | 7.026 | 6.308 | 6.994 |
| $2^{3}D_{2}$ | 2 | 6.787 | 7.399 | 6.852 | 7.403 |
| $2^{3}D_{3}$ | 3 | 6.829 | 7.405 | 6.975 | 7.399 |
| $2^{3}D_{1}$ | 1 | 6.742 | 7.392 | 6.759 | 7.403 |
| $2^{1}D_{2}$ | 2-+ | 6.856 | 7.400 | 6.861 | 7.401 |
| $3^{3}D_{2}$ | 2 | 7.220 | 7.741 | 7.29 | 7.764 |
| $3^{3}D_{3}$ | 3 | 7.256 | 7.750 | 7.409 | 7.764 |
| $3^{3}D_{1}$ | 1 | 7.181 | 7.732 | 7.205 | 7.762 |
| $3^{1}D_{2}$ | 2-+ | 7.296 | 7.743 | 7.302 | 7.762 |
| 1 ³ F ₃ | 3++ | 6.244 | 7.269 | 6.326 | 7.242 |
| 1 ³ F4 | 4++ | 6.312 | 7.277 | 6.501 | 7.244 |
| $1^{3}F_{2}$ | 2++ | 6.175 | 7.273 | 6.182 | 7.234 |
| l^1F_3 | 3+- | 6.307 | 7.268 | 6.335 | 7.241 |
| $2^{3}F_{3}$ | 3++ | 6.801 | 7.616 | 6.876 | 7.615 |
| $2^{3}F_{4}$ | 4++ | 6.859 | 7.617 | 7.041 | 7.617 |
| $2^{3}F_{2}$ | 2++ | 6.740 | 7.618 | 6.741 | 7.607 |
| 2 ¹ F ₃ | 3+- | 6.870 | 7.615 | 6.885 | 7.614 |
| 1^3G_4 | 4 | 6.265 | 7.489 | | |
| 1 ³ G ₅ | 5 | 6.353 | 7.482 | | |
| $1^{3}G_{3}$ | 3 | 6.178 | 7.497 | | |
| l^1G_4 | 4-+ | 6.328 | 7.487 | | |

increases with the increase in the radial quantum number and were found to be in good agreement with the results in previous studies [11,16,24]. It is important to state that the accuracy of the mass spectra is dependent on obtaining a good fit for the phenomenological potential function. This allows for the adjustment of the potential parameters such that the predictions correspond as good as possible to other theoretic works and observed data.

4 Conclusions

The energy bound-state solution of the SSE has been obtained under a spin-dependent Cornell potential energy function using the WKB approximation approach. The energy equation was then used to obtain the mass spectra for the heavy mesons. The mass spectra were found to be in good agreement with the ones obtained by other methods. The charmonium masses deviated from the experimental values with a total percentage error of 3.32%. Also, the bottomonium masses deviated by 1.11%, which indicates an improvement compared to the results in the study by Mansour Gamal [18] and comparable to the works reported in the existing literature [11,24]. The errors were found to reduce significantly for the bottomonium meson due to its heavier mass. The points at which the fitting was carried out may be responsible for the low accuracy of the masses for the low-lying quantum states. However, the SSE together with the phenomenological spin-dependent Cornell potential provides an adequate account of the mass spectroscopy for the heavy mesons and may be used to predict other spectroscopic parameters if the wave function can be obtained.

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References

- [1] Thomson JJ. Cathode rays. Phil Mag. 1897;44:293-16.
- [2] Rutherford E. The scattering of α and β particles by matter and the structure of the atom. Phil Mag. 1911;21(125):669–88.
- [3] Ryder LH. Quantum field theory. Cambridge: Cambridge University Press; 1985.
- [4] Mann R. An introduction to particle physics and the standard model. Boca Raton: CRC Press; 2021.
- [5] Li BQ, Meng C, Chao KT. Coupled-channel and screening effects in charmonium spectrum. Phys Rev D. 2009;80:014012.
- [6] Pininnington MR, Wilson DJ. Decay channels and charmonium mass shifts. Phys Rev D. 2007;76:077502.
- [7] Mutuk H. S-wave heavy quarkonium spectra: Mass, decays, and transitions. Adv High Energy Phys. 2018;2018:5961031.
- [8] Ali MS, Yasser AM, Hassan GS, Moustakidis CC. Spectra of quarkantiquark bound states *via* two derived QCD potentials. Quant Phys Lett. 2016;5(1):7–14.
- [9] Ali MS, Hassan GS, Abdelmonem AM, Elshamndy SK, Elmasry F, Yasser AM. The spectrum of charmed quarkonium in non-relativistic quark model using matrix Numerov's method. J Radiat Res Appl Sci. 2020;13(1):226–33.
- [10] Chaturvedi R, Rai AK. B_c meson spectroscopy motivated by general features of pNRQCD. Eur Phys J A. 2022;58:228.
- [11] Mansour H, Gamal A, Abolmahassen M. Spin splitting spectroscopy of heavy quark and antiquarks systems. Adv High Energy Phys. 2020;2020:1–11.
- [12] Bhavsar T, Shah M, Vinodkumar PC. Status of quarkonia-like negative and positive parity states in a relativistic confinement scheme. Eur Phys J C. 2018;78:227.
- [13] Kher V, Rai AK. Spectroscopy and decay properties of charmonium. Chin Phys C. 2018;42(8):083101.
- [14] Deng WJ, Liu H, Gui LC, Zhong XH. Spectrum and electromagnetic transitions of bottomonium. Phys Rev D. 2017;95:074002.
- [15] Gupta P, Mehrotra I. Study of heavy quarkonium with energy dependent potential. J Mod Phys. 2012;3:1530–36.
- [16] Ebert D, Faustov RN, Galkin VO. Spectroscopy and Regge trajectories of heavy quarkonia and mesons. Eur Phys J C. 2011;71:1825.

- [18] Mansour H, Gamal A. Meson spectra using Nikiforov-Uvarov method. Result Phys. 2022;33:105203.
- [19] Boroun GR, Abdolmalki H. Variational and exact solutions of the wavefunction at origin (WFO) for heavy quarkonium by using aglobal potential. Phys Scr. 2009;80:065003.
- [20] Chen H, Zhang J, Dong YB, Shen ON. Heavy quarkonium spectra in a quark potential model. Chin Phys Lett. 2001;18(12):1558.
- [21] Barnes T, Godfrey S, Swanson ES. Higher charmonia. Phys Rev D. 2005;72:054026.
- [22] Cao L, Yang YC, Chen H. Charmonium States in QCD-inspired quark potential model using Gaussian expansion method. Few-Body Syst. 2012;53:327–42.
- [23] Kher V, Chaturvedi R, Devlani N, Rai AK. Bottomonium spectroscopy using Coulomb plus linear (Cornell) potential. Eur Phys J Plus. 2022;137:357.
- [24] Soni NR, Johi BR, Shah RP, Chauhan HR, Pandya JN. $Q\bar{Q}(Q \in \{b, c\})$ spectroscopy using the Cornell potential. Eur Phys J C. 2018;78:592.
- [25] Maireche A. A new model to describe Quarkonium systems under modified Cornell potential at finite temperature in pNRQCD. Int J Phys Chem. 2022;88:1–16.
- [26] Maireche A. The relativistic and nonrelativistic solutions for the modified unequal mixture of scalar and time-like vector Cornell potentials in the symmetries of noncommutative quantum mechanics. Jordan J Phys. 2021;14(1):59–70.
- [27] Maireche A, Imane D. A new nonrelativistic investigation for spectra of heavy quarkonia with modified Cornell potential: Noncommutative three dimensional space and phase space solutions. J Nano- Electron Phys. 2016;8(3):03024.
- [28] Omugbe E, Aniezi JN, Inyang EP, Njoku IJ, Onate CA, Eyube ES, et al. Non-relativistic mass spectra splitting of heavy mesons under the Cornell potential perturbed by Spin–Spin, Spin–Orbit and tensor components. Few Body Syst. 2023;64:66.
- [29] Ansler C, Doser M, Antonelli M, Anser DM, Babu KS, Baer H, et al. Particle data group. Phys Lett B. 2008;667:010001.
- [30] Workman RL, Burkert VD, Crede V, Klempt E, Thoma U, Tiator L, et al. Particle data group. ProgTheor Exp Phys. 2022;2022:083C01.
- [31] Patrignani C, Agashe K, Aielli G, Amsler C, Antonelli M, Asner DM, et al. Particle data group. Chin Phys C. 2016;40:100001.
- [32] Omugbe E, Osafile OE, Okon IB, Inyang EP, William ES, Jahanshir A. Any *l*-state energy of the spinless Salpeter equation under the Cornell potential by the WKB approximation method: An application to mass spectra of mesons. Few Body Syst. 2022;63:6.
- [33] Fulcher LP, Chen Z, Yeong JC. Energies of quark-antiquark systems, the Cornell potential, and the spinless Salpeter equation. Phys Rev D. 1993;47(9):4122–32.
- [34] Salpeter EE. Mass corrections to the fine structure of hydrogen-like atoms. Phys Rev. 1952;87(2):328–43.
- [35] Zarrinkamar S. Quasi-exact solutions for generalised interquark interactions in a two-body semi-relativistic framework. Z Naturforsch. 2016;71(11):1027–30.
- [36] Langer RE. On the connection formulas and the solutions of the wave equation. Phys Rev. 1937;51:669–76.